



## MELOS: Japan's Mars Exploration Plan

Exploration with

Orbiter

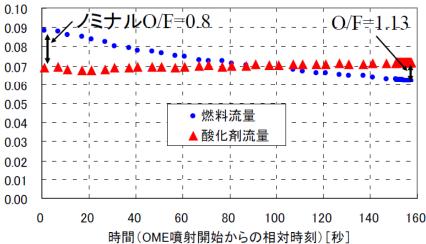
Synergy

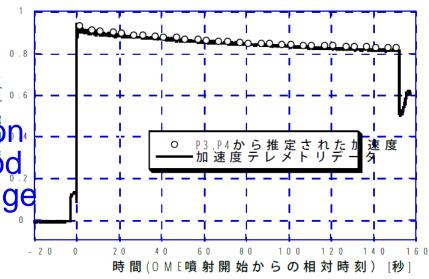
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- VOI attempted on 7 Dec 2010 but failed...
  - Likely cause is "clogged" check <sup>0.0</sup> valve in the fuel line (CV-F)
    - It prevented smooth flow of fuel, resulting in "less fuel than oxidizer" of condition (O/F > 1).
      - The excess fuel acts as "coolant" for the thruster throat & nozzle. This does *not* work if the fuel is less than the oxidizer.
    - Without enough cooling, a damage has occured to the thruster, and the spacecraft went to "Safe Hold".
  - The spacecraft (including mission instruments) seems to be in good condition, and we will re-challenge VOI 6 years later.





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lar	Launch	Arrival	Mass to Mars Orbit (H-IIA202)	Mass to Mars Orbit (H-IIA204)
	Jan 2019	Feb 2022	1.4t	2.4t
Un	Jul 2020	Feb 2021	1.2t	2.1t
C. //	Nov 2021	Jan 2024	1.3t	2.2t
J	Sep 2022	Aug 2023	1.1t	2.0t

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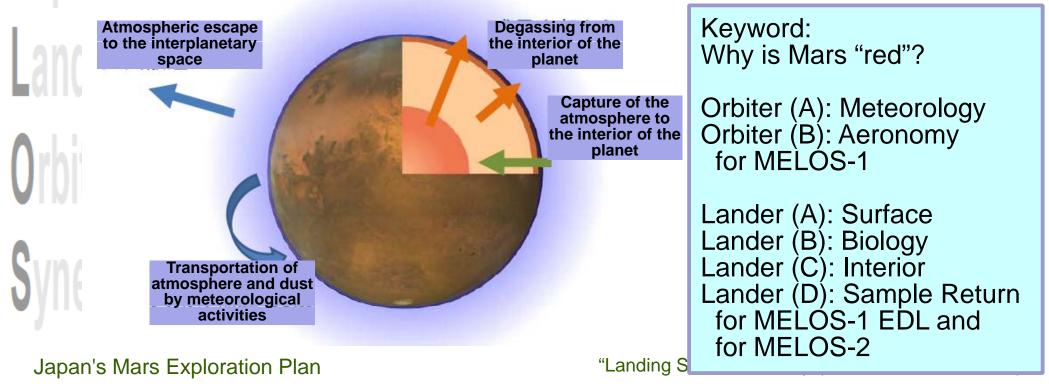




Understanding the Martian System

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- Interior + surface + atmosphere + surrounding space
  - To understand the evolution and to answer the fundamental question "Why (and how) is Mars different from the Earth?", missions designed to study inter-relations between these are needed.
  - Both "orbiting" science and "landing" science are important.



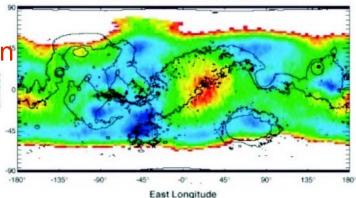


- Comparative Meteorology (Earth vs Venus vs Mars)
  - Similarity: rotation period, tilt of the pole
    - Tenuous CO<sub>2</sub> atmosphere
       vs suspended dust (heat source)
    - Large seasonal variation (eccentricity)
       vs relatively small thermal inertia
    - Episodic "global" dust storm
    - Underground water (ice) reservoir
- Transportation/Relocation of Water & Dust
  - Never been studied in detail

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- Limitation of "local-time fixed" orbit
- Need to characterize "global" transportation
- Especially in the lower-most atmosphere
- 3-D structure of temperature, composition, isotopic ratio, etc.

Water Abundance (weight percent)

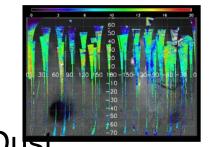


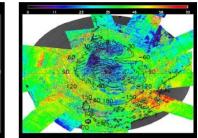
"Landing Sites" Workshop (20 Jan 2011 @ ESTEC)

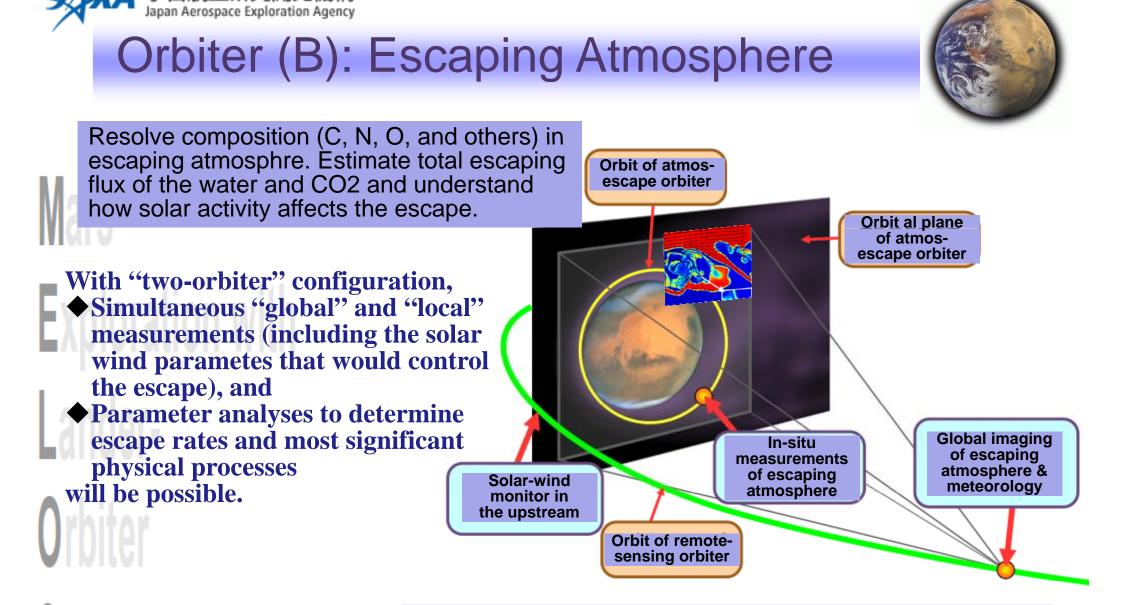


tilt of the pole









Proposed Payload

Orbiter:

宇宙航空研究開発機構

"Escaping Atmosphere" lon analyzers (mass, energy, velocity), Neutral-gas mass analyser, Langmure probe, Magnetometer, Electric-field & Plasma-wave package, Potential control

"Remote Sensing" **Orbiter:** 

Global imager for "escaping atmosphere", UV absorption cell, Solar-wind monitor. Solar-radiation monitor

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## Lander (A): Surface Environment



- Goal: Understand climate changes of 100s My scale and chemical coupling with thermochemical evolution of the solid planet.
- Objectives:
- (a) Discover initial melt through composition analysis of basaltic flow to infer composition of the mantle.
  - (b) Understand ancient climate by investigating sedimentary rocks.
- Method: Geological survey over a great distance with a well-equiped rover.
- Candidate sites for survey (landing):
  - (a) Ejecta of a young crator (~10 km dia.) in basaltic region.
     (b) Nili Fossae, or ancient coast line.

### • Model Payload:

3-D camera with variable LCF, Macro spectroscopic camera, XRFD, LIBS, K-Ar Dating, Magnetometer, Radar sensing, Electro-magnetic field, Meteorology package

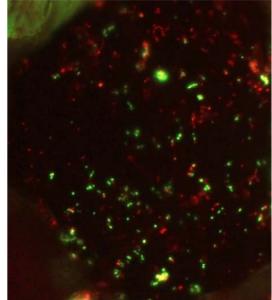
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## Lander (B): Astrobiology

#### Most likey place for Martian life?

- Surface soil near the methane vent
  - Terrestial life may survive at a few cm below Martian surface.
  - Martian environment ~4 Gy ago similar to Earth.
  - The birth of life can be in very short time.
  - Martian methane detected recently.
  - Bacteria that utilize methane & iron oxide discovered.
- Method?
  - With fluorochrome and a microscope
    - Dye protein, membrane, catalyst, etc.
    - Target sensitivity: 10 cells / 1 g of soil. (compared to 10<sup>4</sup> cells / 1 g in Earth desert)
    - Conclude "no life on Mars" if no detection.
    - Possibly detect organic materials related to chemical evolution before life.



Antarctic soil after dying







## Lander (C): Interior Structure

Rotation

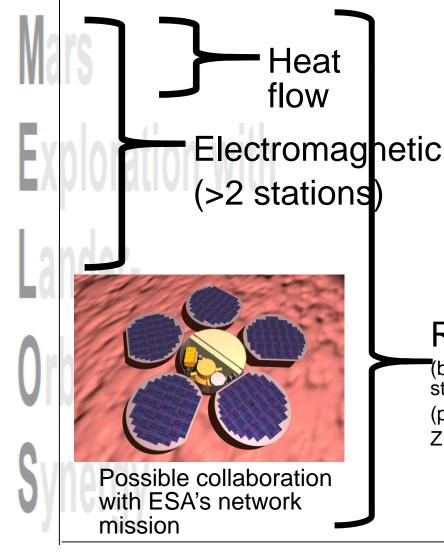
Z Term of Mars

stations)

(better to have multiple

(precession, nutation)

(Depth)



# ure Preliminary Reference Interior Structure Model on Mars (PRISM<sup>2</sup>)

Geo-Electromagnetic

Seismic

#### Mars-quake

Geodetic

(better achievement with multiple stations) (guakes, free oscillations)

#### (Information)

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Lander (D): No-Landing Sampel Return (MASC)



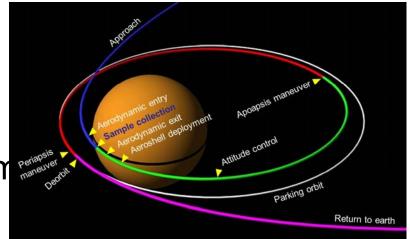
Capture and return the dust & atmospheric samples by performing aero-flyby.

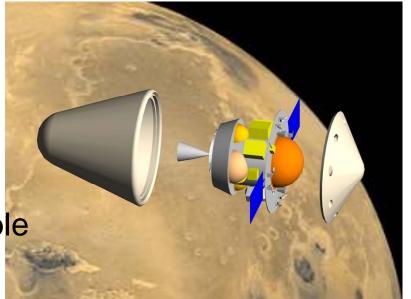
 – "Global average" information from such sample is significant to:

- Orbiter sciences (meteorology, atmospheric escape),
- Biological study

 Benefit from state-of-the-art analyses in the laboratory

> High-sensitivity & high-precision analyses that would not be possible with *in-situ* equipment







#### Technology Development: Surface Exploration with 100-kg Class Rover



Robotics for geological & biological survey
Rover (mobility, obstacles, durability against dust)
Manipulator (function, portability, etc.)

 Under development by a team of people from JAXA and from many universities







Technology Development: Aerodynamic Control



## Aerobraking

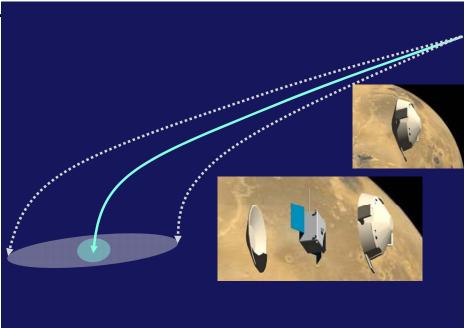
 HITEN (Engineering Demonstration Mission): Multiple swingbys between the Moon and the Earth + Aerobraking experiments at the Earth

## Atmospheric Entry/Landing

- HAYABUSA sample-return capsule
  - SELENE-2 Lunar Lander (under development)

## Aero-assisted Control

Pin-point landing, non-landing sample return (LANDER D)



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