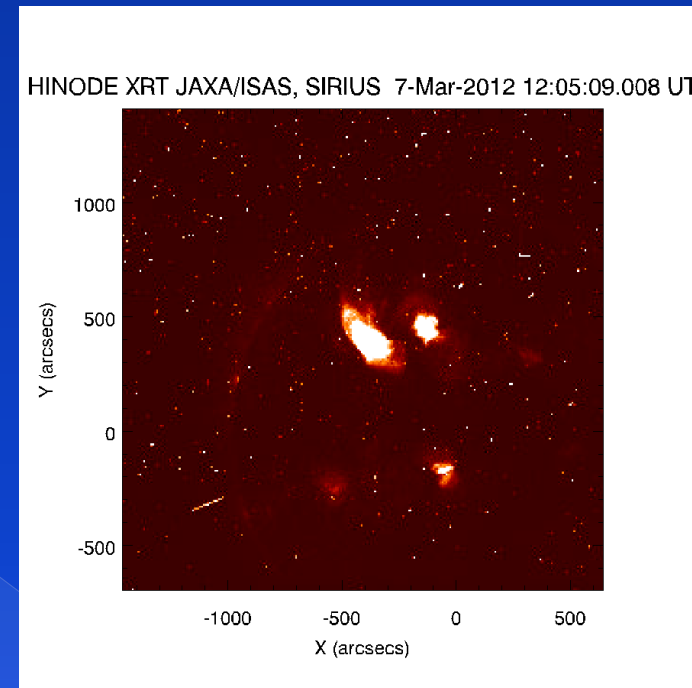


Fluctuating the Speckle is detected by 'Hinode' XRT according to solar activity

Masanori Yamada
Ibaraki University M2

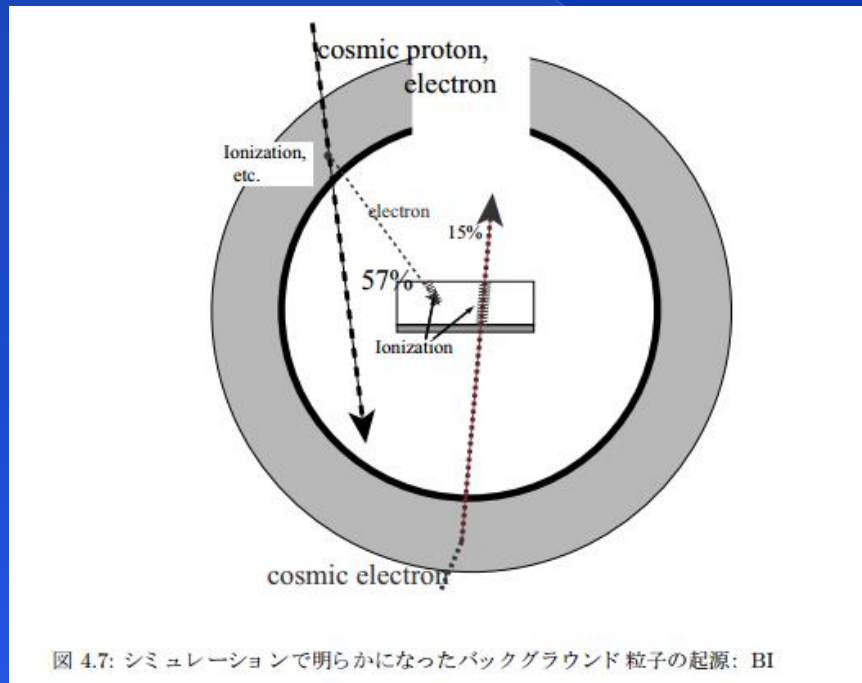
Solar Energetic Particle and Space Weather

- Flare and CME generate SEPs.
 - SEPs generate “false” signal on the CCD image.
 - This study calls Speckle.
- Cause is secondary particles generated by collision with shield of SEPs.
 - Speckle becomes indication of how many SEPs flew on the satellite orbit .



Hinode / XRT
2012 / 03 / 07 UT

Cause of speckle



Quote from Master thesis
Katou 2007

- I. High energy particle collisions to shield.
- II. Electron generation occurs.
- III. The collision in CCD.
- I. High energy particle collisions in CCD.

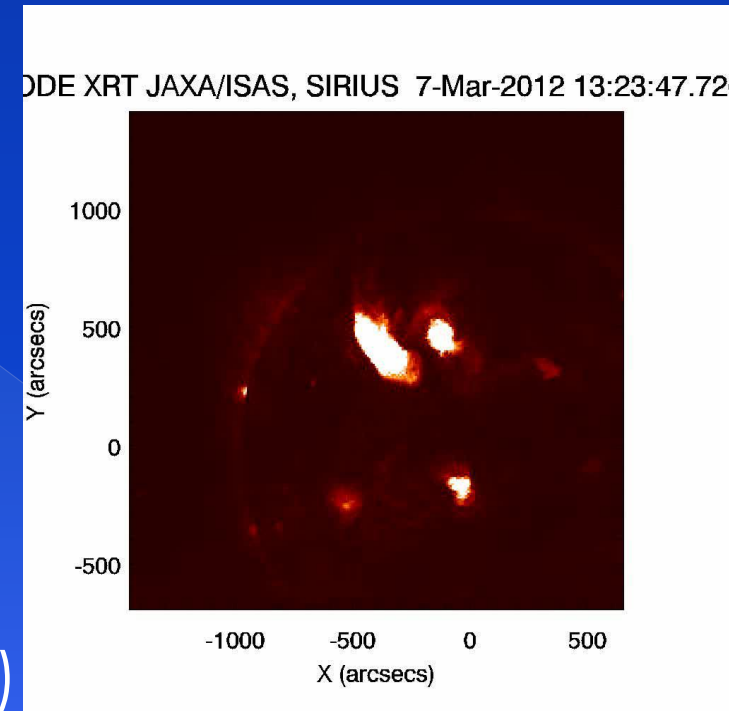
Purpose

- SEP will cause a lot of space weather damage.
 - However, quantitative model of the SEP is not complete .
- For the character clarification of the SEP , it will focus on the CCD noise .
- Performs detection analysis of the Speckle using satellite data, to identify the cause of the variation.
 - It will continue to clarify the conditions to be observed .

Observation equipment

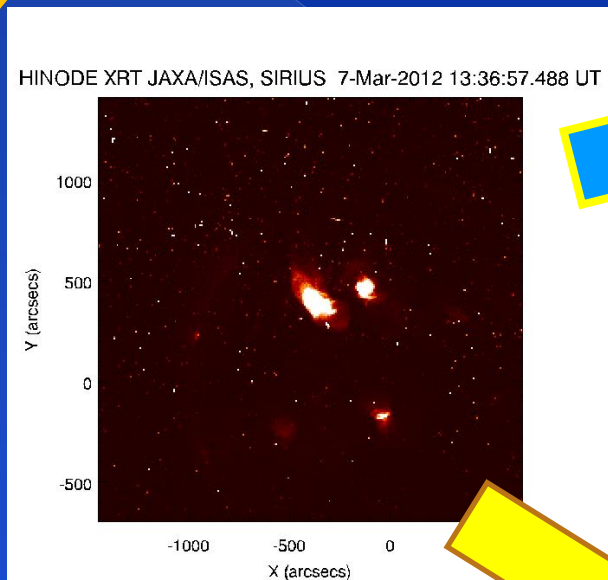
Using data

- Used equipment
 - Hinode / X – Ray Telescope (XRT)
- Orbit
 - Sun-synchronous polar orbit
 - Altitude 680 km
- Used data
 - Flare time patrol image
 - Image size : 256×256 pix
 - Time resolution : 30 sec
 - Used filter : Ti_poly
 - Exposure time : 8.8msec(Ave.)
- Analysis object
 - Solar proton event



Hinode / XRT
7th Mar. 2012

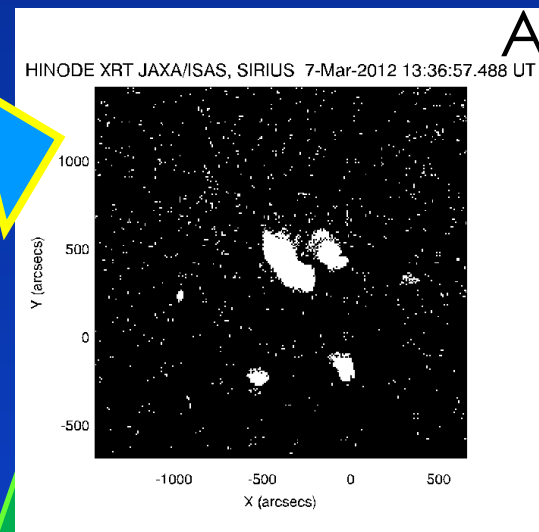
Image processing



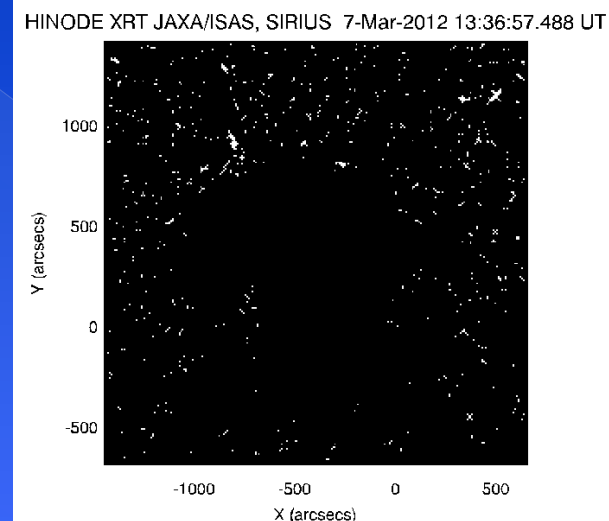
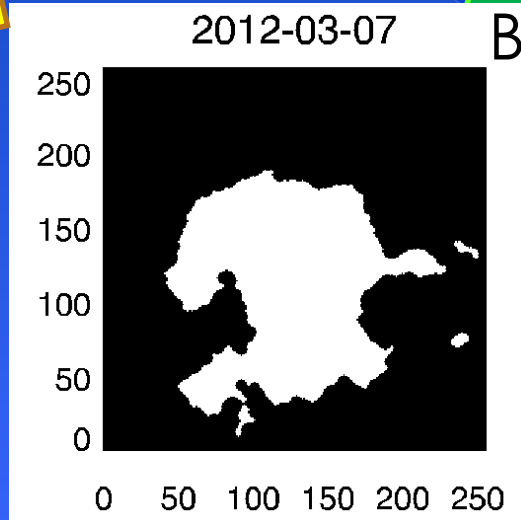
Emphasize speckles
of row data.

Binarization of
superimposed images.

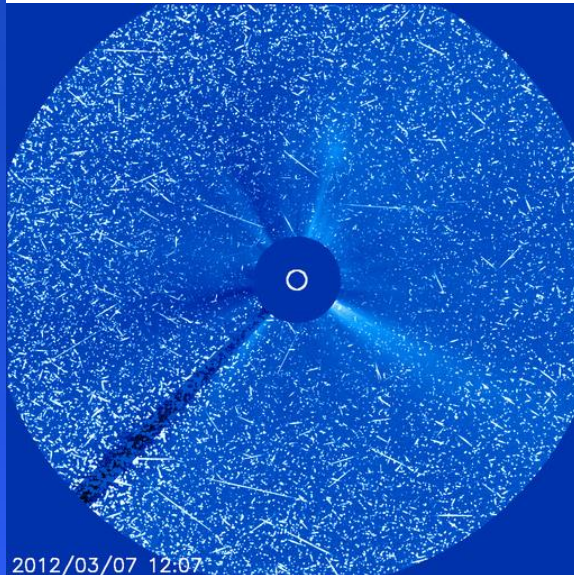
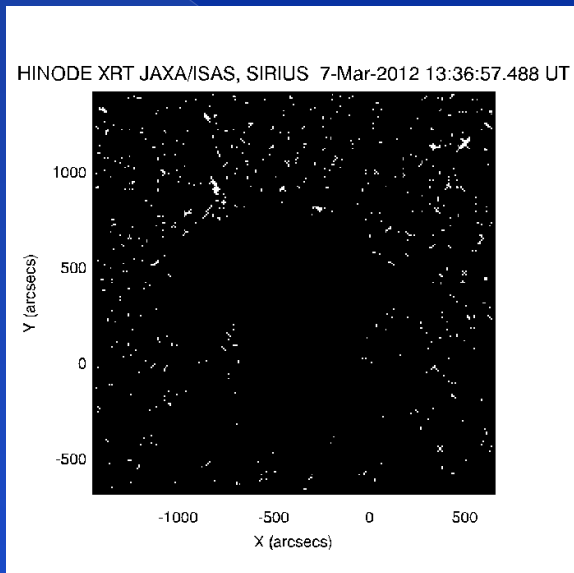
Binarization of
single image



Subtract B from A



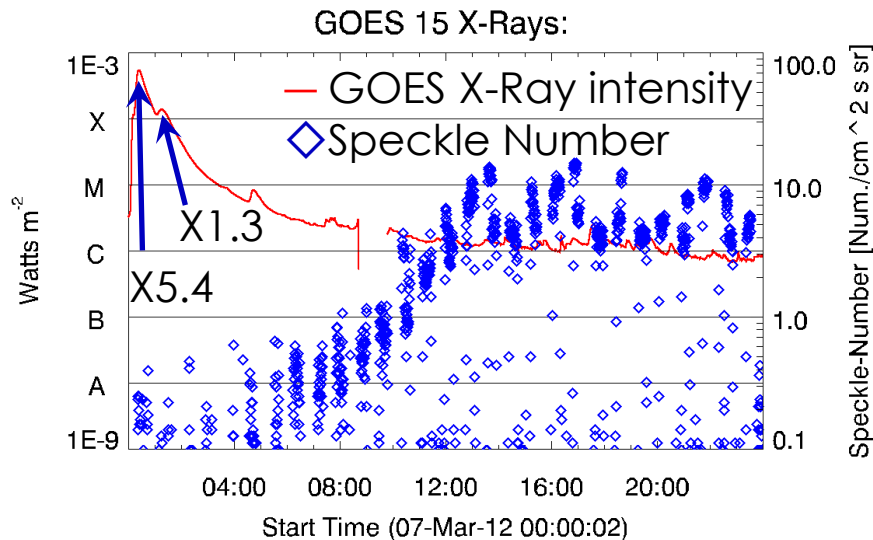
Character of speckle in Hinode / XRT



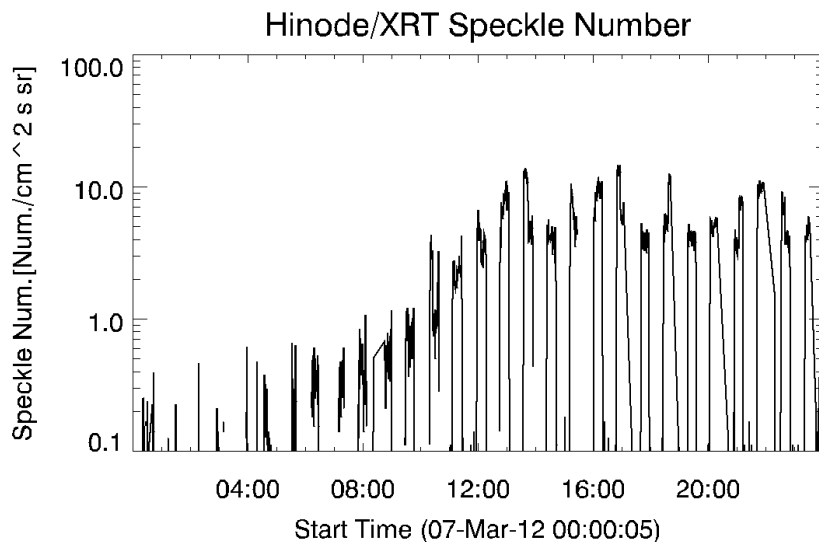
SOHO / LASCO (C3)
7th Mar. 2012

- Speckle size is small and short than SOHO/LASCO's it .
- CCD camera of the XRT is made Si.
 - The energy required to produce the Speckle could be estimated to about 50keV per pixel .
 - I continue calculated the energy of the particles at the time of collision.

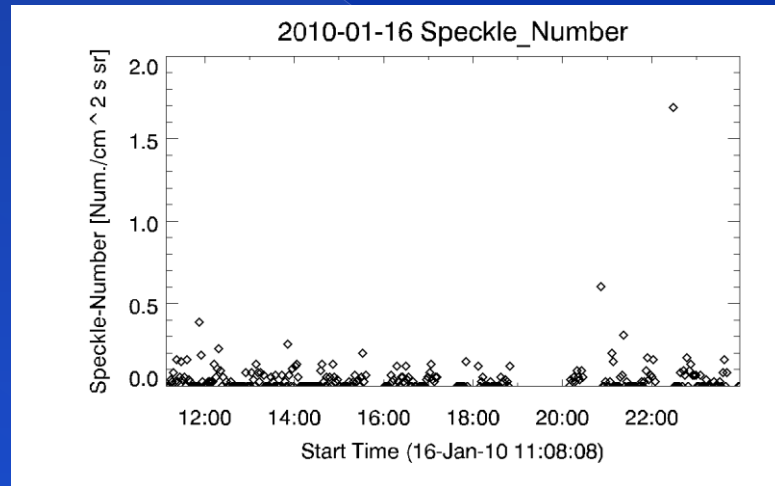
Time variation of detected number of speckles



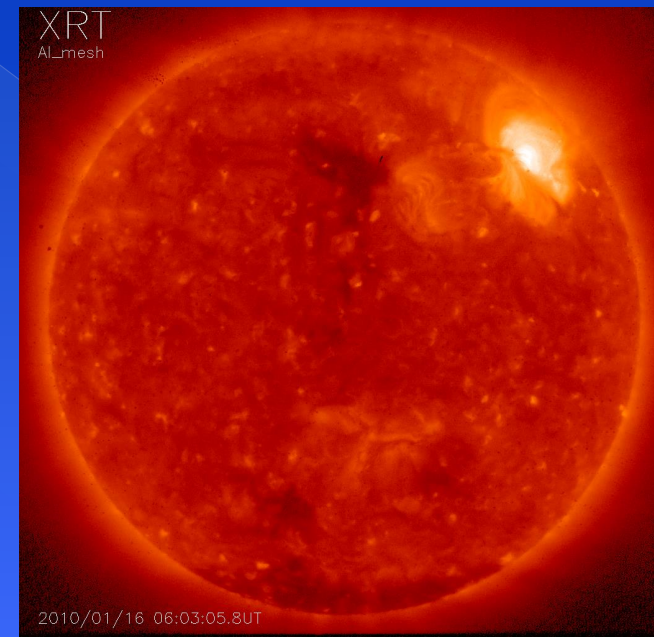
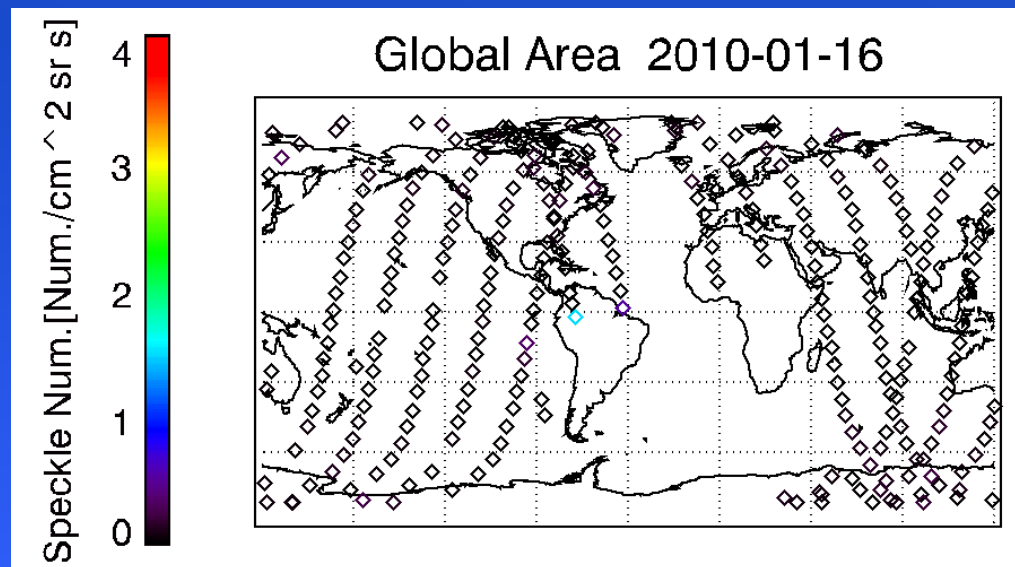
- Analysis results of 7th Mar. 2012.
- X5.4 and X1.3 Flare occurred 0:02 and 1:05 UT.
- After flares occur , speckles increase after a few hours.
- There is a cycle in the increase or decrease of the speckles.



Consideration for background

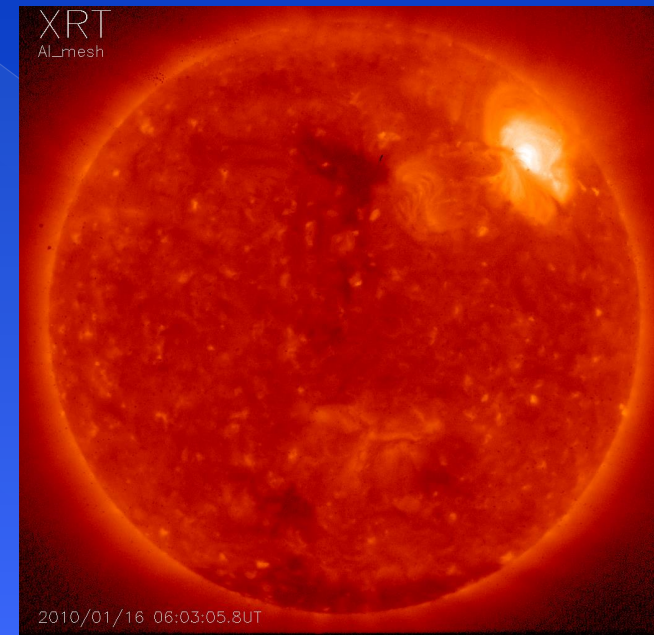
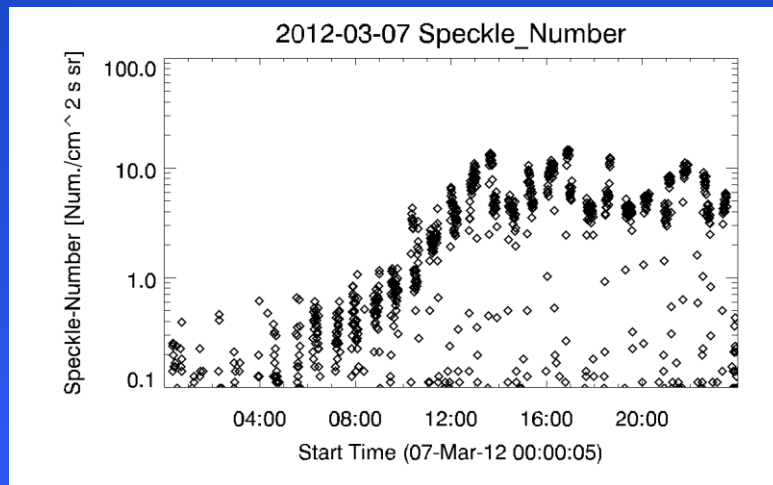
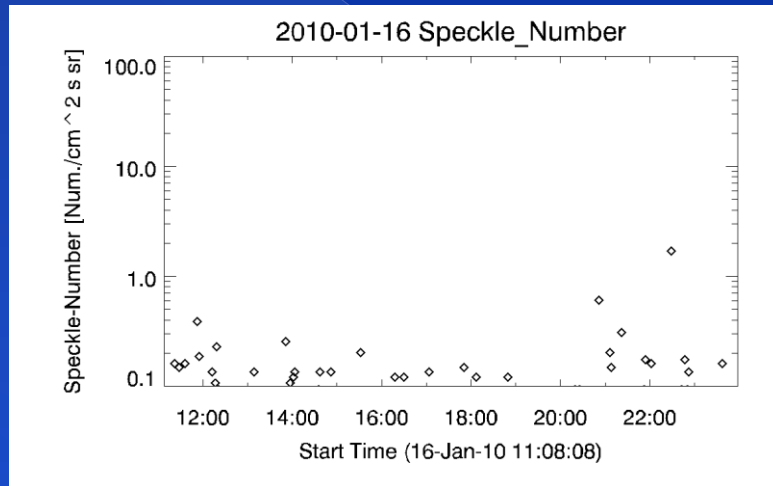


- Analyzed data for 16th Jan. 2010 .
 - AR are less .
 - Large flare and CME didn't occur .
 - $Med.BG_{speckle} = 0.013$

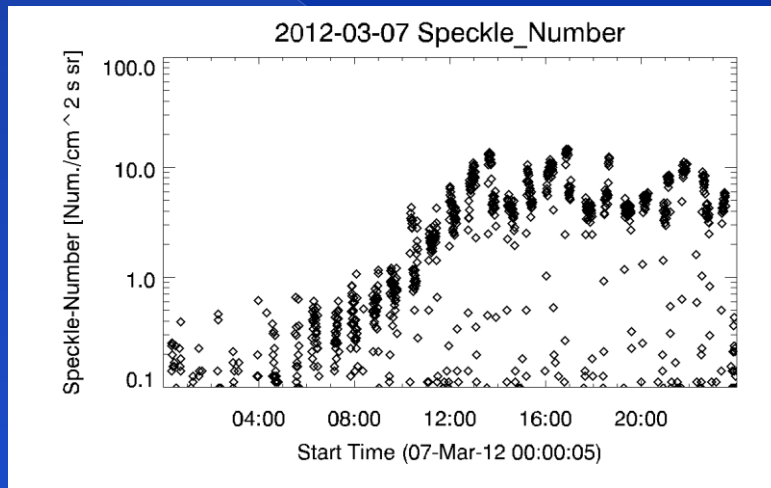


Consideration for background

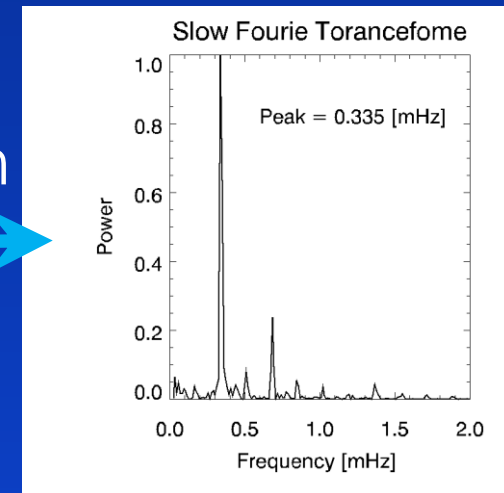
- Analyzed data for 16th Jan. 2010 .
 - > AR are less .
 - > Large flare and CME didn't occur .
 - > $Med.BG_{speckle} = 0.013$



Periodicity of speckle number



Fourier transform



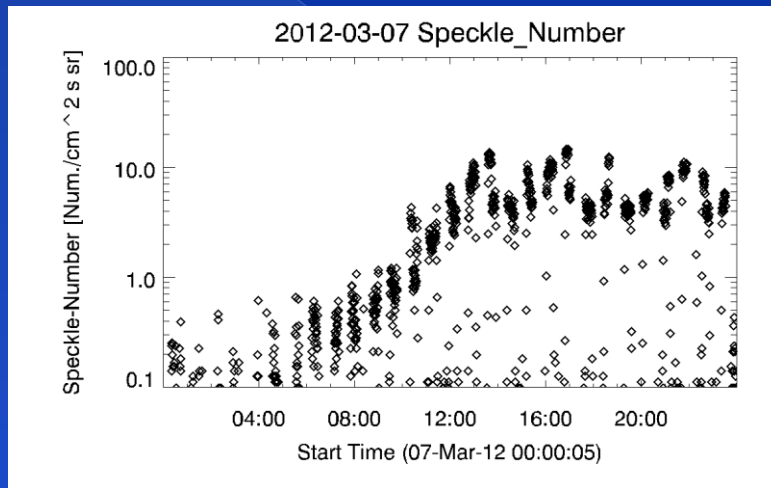
- Detection number increased in certain timing .

- There is an increase of decrease in a particular location .

Hinode
Orbital period 5900 [sec]

Speckle
Detection cycle 2980 [sec]
F = 0.335

Periodicity of speckle number

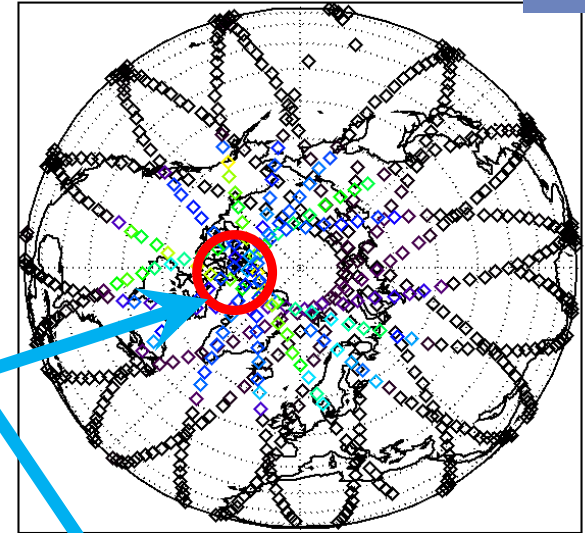


- Speckle has been found to increase in the polar region in orbit .
- It increased at 60 degrees or more from latitude 55 degrees is happening .

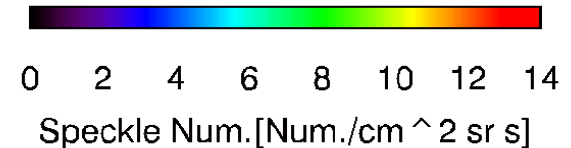
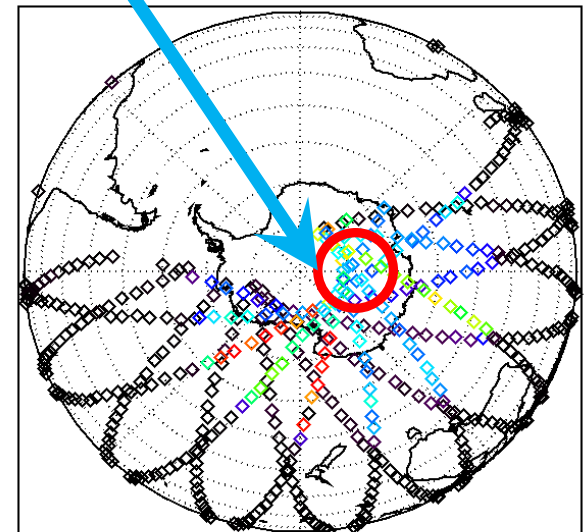
Magnetic pole

Nor. Hem. 2012-03-07

12

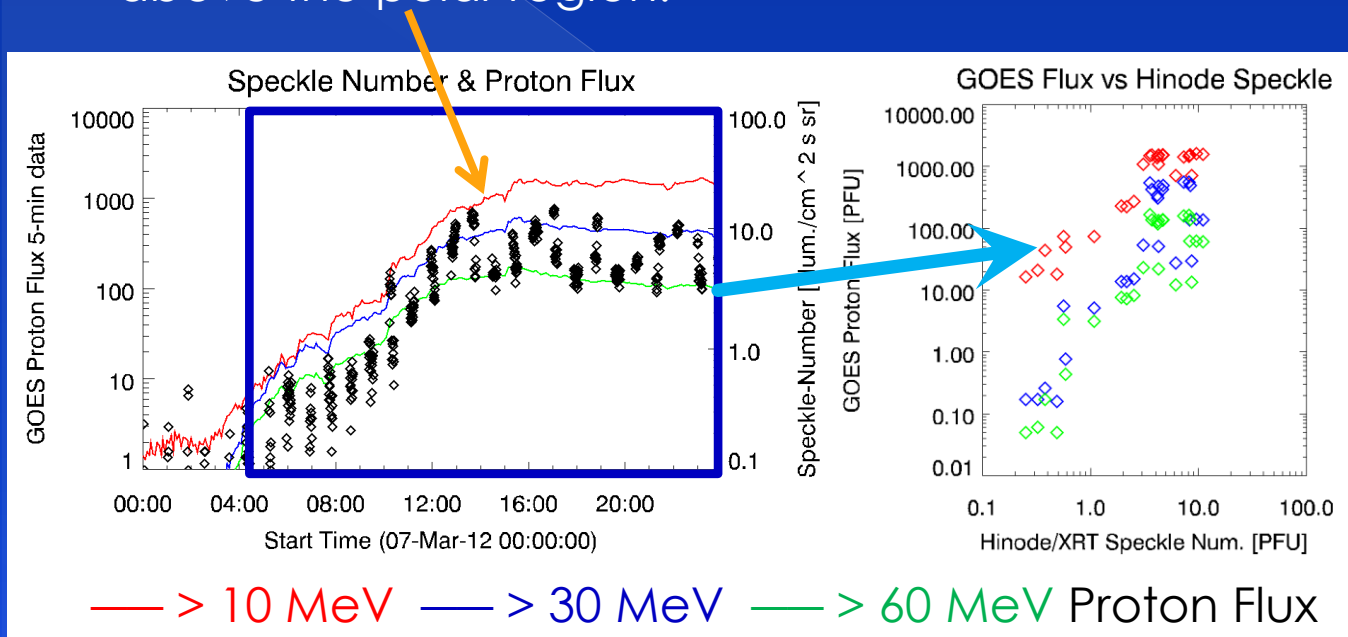


Sou. Hem. 2012-03-07



Comparison GOES Proton Flux

During the Hinode moving
above the polar region.



X5.4 UT 0:02

X1.3 UT 1:05

2012 / 03 / 07 Correlation
factor

> 10 MeV 0.92

> 30 MeV 0.89

> 60 MeV 0.90

- Speckles were a change similar to the GOES .
 - Using the data of a plurality of satellites, we will consider the effects of the surrounding environment .

Summary

- Speckle is an index to measure the inflow of particles.
- Using the data of Hinode / XRT, it was detected Speckle.
- Speckles were a change similar to the GOES above the polar region.
- The energy required to produce the speckle was estimated to 50keV per 1pixel.
- Go conducted consider plasma flow path and energy.