# 144th Hinode SSC Meeting on 24th January, 2019 at 07:00 JST

#### Short Summary, Conclusions and Actions

#### a. Program Status

#### 1. Instrument Status Review

**SOT** operating without its Filtergraph (FG) camera following an electronic fault. Spectro-Polarimeter (SP) and Correlation Tracker (CT) are nominal.

**XRT** is nominal.

EIS is nominal.

### 2. Report on Changes to Instrument Telemetry Allocation

There are no further reports on telemetry allocation changes

### 3. FM Calendar

Hinode focus mode calendar has been updated.

## 4. HOP Prioritisation

SSC asked by **SWG** to prioritise HOPs i) with associated ground-based observations that were overlapping in a time zone and ii) in cases that generated mission telemetry use conflicts.

#### ACTION: Culhane to ensure that such cases were highlighted in SSC meeting notes; Ongoing

#### **b.** Previous Action Items

Mckenzie confirmed May 14th as the new EUNIS launch date

**Savage** stated that the list of programmes from the previous Senior Review submission would now be circulated following the end of the partial government closure

#### c. Review/Discussion of Open HOPs and ToOs

- routine HOPs 79, 81 and 130 were run as planned during December
- NU-STAR completed an observation with Hinode and the VLA; good data obtained
- HOP 364 will be run following IRIS eclipse season end in February; De Pontieu to confirm date
- agreed post-meeting to be run during week of 17th February
- HOP 366 will continue on a weekly schedule; Watanabe will update the monthly events list

#### d. Review of New or Updated Proposals and Scheduling of Observations

# 1. Long Period Pulsations of Plasma Velocity and Density in Loops - Pelouze (gabriel.pelouze@ias.u-psud.fr), Auchere, Bocchialini, Parenti, Culhane/Harra/SSC; ToO HOP 367

- to detect and characterise plasma downflows in loop footpoints that are associated with long-period intensity pulsations of ~ 3 - 16 hours
- observations to be made with **EIS**
- target suitable active region and track it for most of its lifetime but for at least 2 days continuously during CM passage; prefer 6 days. Observe for ~ 10 hr/day.
- run study 571 (ar\_vel\_fast\_scan); obtain raster every 40 min
- use spectra to characterise variations of plasma velocity, temperature and density
- run on suitable AR; observing window to start in focussed mode on 29<sup>th</sup> January

Continuing monthly observations are:

- Polar Monitoring Shimojo; CORE HOP 81
- run on 7<sup>th</sup> February (N pole fast) and 9<sup>th</sup> February (S pole fast)
- Synoptic SOT Irradiance Scans Tarbell; CORE HOP 79
- run on 14<sup>4h</sup> February (N/S only)
- Multi-temperature Full Disk Slot Scans Ugarte-Urra, Brooks, Warren; CORE HOP 130
- run on 2<sup>nd</sup> and 26<sup>th</sup> February
- Cycle 25 Bright Points Bryans; HOP 336
- run on every Monday when feasible

## e. Monthly Science Reports

- next **Hinode** monthly science report will be prepared by the **SOT Team** for 8<sup>th</sup> February
- **NOTE**: Science chart site access has been changed due to IT requirements; **Savage** has established a new Google drive site for template and previous chart
- provide one summary slide for Hinode team management at MSFC and two additional slides for NASA HQ

## f. Date of Next Meeting

next meeting: 21<sup>st</sup> February, 2019 at 07:00 JST; 20<sup>th</sup> February, 2019 as appropriate in US/Europe

## g. AOB

**Savage** reported that she and **Elrod** will monitor operations and address critical issues only until the US Government partial shutdown has ended.